### Looking forward to new physics and neutrinos at the LHC

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### High Energy, Cosmology and Astro-particle Physics (HECA) May 25, 2021

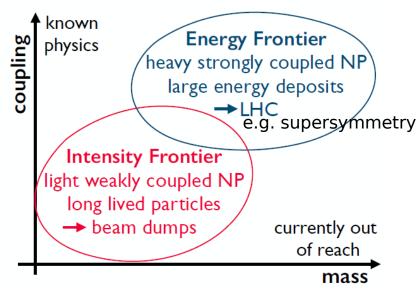


## OUTLINE

- Motivation & Fundamentals
- Far-forward BSM physics at the LHC
  - new physics production in the far-forward region of the LHC
  - selected BSM models
  - BSM particle production away from the ATLAS IP,
  - light dark matter (DM)
- High-Energy neutrino physics at the LHC
- Additional opportunites (QCD, connections to cosmic ray physics,...)
- Concluding remarks

# MOTIVATION

#### LIGHT NEW PHYSICS

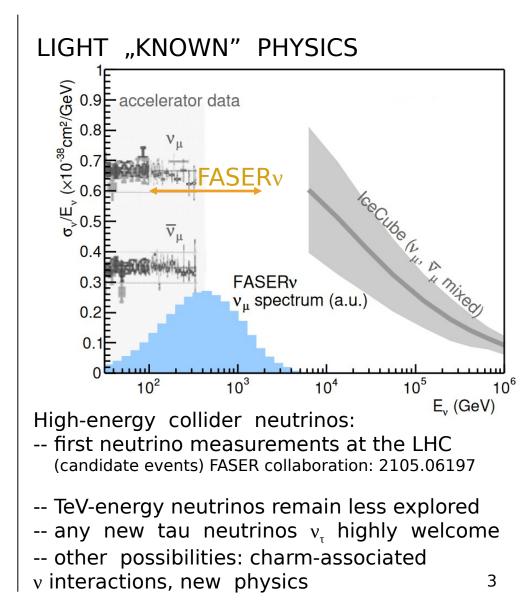


-- "leave no stone unturned"

-- cosmology (dark matter, inflation, bariogenesis,...)

-- neutrino masses (GeV-scale heavy neutral leptons)

-- anomalies



# (SELECTED) BSM CONNECTIONS

a) gauging global symmetries of the SM e.g.  $U(1)_{Le-L\mu}$ ,  $U(1)_{B-L}$ , i = i new dark vector

M. Bauer, P. Foldenauer, J. Jaeckel, **JHEP 1807 (2018) 094** 

Additional U(1) groups might arise in extensions of the SM group,

typically light new gauge boson must be very weakly coupled to the SM

Mohapatra R N and Senjanovic G, Phys. Rev. D23:165 (1981

Kinetic mixing between the photon and new vector can also be loop-induced in secluded regime  $U(1)_{v}$ 

b) mirror sector / Twin Higgs scenarios often predict new scalars coupled to the SM via Higgs

 $(H^{\dagger}H) \times m_{H}^{2} \longrightarrow (H^{\dagger}H) \times (m_{H}^{2} + c_{1}S + c_{2}S^{2} + \ldots),$ 

Lanfranchi etal, 2011.02157

c) Right-handed neutrinos e.g. vMSM (v masses and oscillations, DM, baryon asymmetry) T. Asaka, S. Blanchet and M. Shaposhnikov, *Phys. Lett.* **B631** (2005) 151-156 T. Asaka and M. Shaposhnikov, *Phys. Lett.* **B620** (2005) 17-26

d) Axion is an example of light weakly coupled particle postulated long time ago
 later generalized to axion-like particles

### IDDEN SECTOR PORTALS

- new ",hidden" particles are SM singlets (but gauged  $U(1)_{B-L}$  etc. are also considered) - interactions between the SM and "hidden" sector arise due to mixing through some SM portal

$$\mathcal{L}_{\text{portal}} = \sum O_{\text{SM}} \times O_{\text{DS}}$$

B. Patt, F. Wilczek, 0605188 B. Batell, M. Pospelov, A. Ritz, 0906.5614

**Renormalizable** Coupling Portal Dark Photon,  $A_{\mu}$  $-\frac{\epsilon}{2\cos\theta_W}F'_{\mu\nu}B^{\mu\nu}$  $(\mu S + \lambda S^2) H^{\dagger} H$ Dark Higgs, S $\frac{a}{f_a}F_{\mu\nu}\tilde{F}^{\mu\nu}, \ \frac{a}{f_a}G_{i,\mu\nu}\tilde{G}_i^{\mu\nu}, \ \frac{\partial_{\mu}a}{f_a}\overline{\psi}\gamma^{\mu}\gamma^5\psi$ Axion, a Sterile Neutrino, N $y_N LHN$ PBC report, 1901.09966

# FAR-FORWARD BSM PROGRAM

**IDEA** 

#### Forward BSM & neutrino physics at the

J.L. Feng, I. Galon, F. Kling, ST, 1708.09389 FASER Collaboration: 1811:10243, 1812.09139 1908.02310, 2001.03073

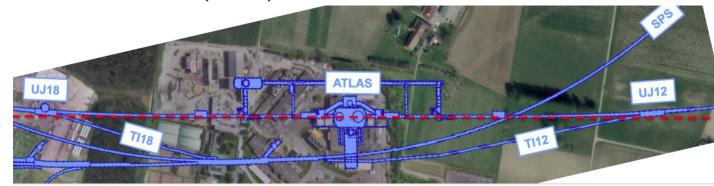
### ForwArd Search ExpeRiment (FASER) – small (~0.05 m<sup>3</sup>) and detector to be placed few hundred meters downstream the ATLAS IP to harness large, currently "wasted" forward LHC cross section $\sigma_{inel} \sim 75 \text{ mb, e.g., } N_{\pi} \sim 10^{17} \text{ at } 3/ab^{-1} (for comparison } \sigma \sim fb - pb, e.g., N_{H} \sim 10^{7} at 300 \text{ fb}^{-1} in high-p_{T} searches)$ new FASER physics & neutrinos VERY SCHEMATICALLY (side tunne p collision axis ATLAS IP FASER 7

#### Forward BSM & neutrino physics at the LHC

FASER LoI & TP: 1811:10243, 1812.09139 FASERv LoI & TP: 1908.02310, 2001.03073

# RUN 3 & HL-LHC PLANS

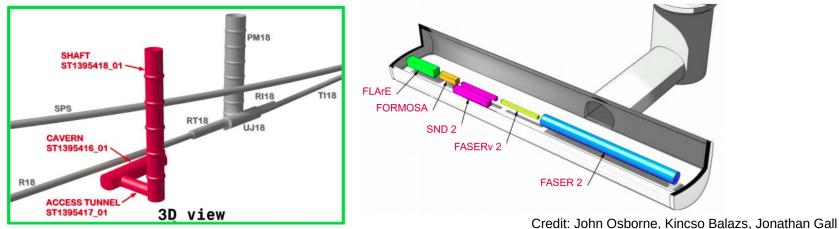
#### **Run 3 main FASER -- cylindrical detector:** $L = 1.5 \text{ m}, R = 10 \text{ cm}, V = 0.05 \text{ m}^3, 150 \text{ fb}^{-1}$ (Run 3)



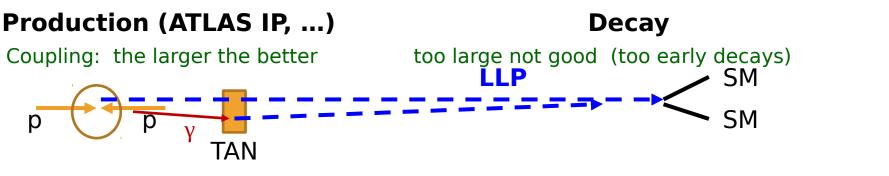
#### HL-LHC

(possible upgrade) FASER 2: *L* = 5 m, *R* = 1 m, V = 16 m<sup>3</sup>, 3 ab<sup>-1</sup> (HL-LHC)

### Forward Physics Facility R.M.Abraham etal, Snowmass 2021 LoI



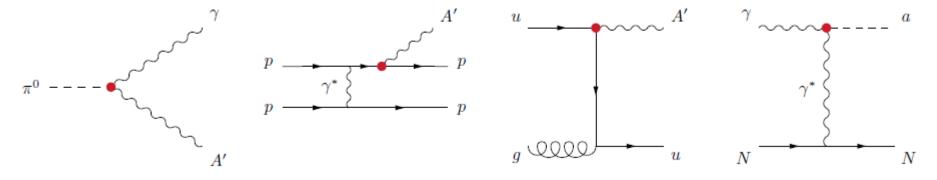
### SEARCH FOR HIGHLY DISPLACED DECAYS



Various production mechanisms: -- meson decays (light & heavy)

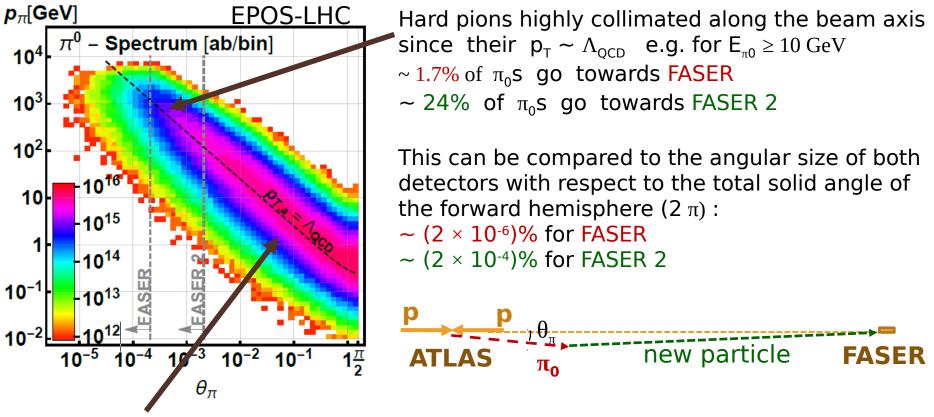
$$N_{\rm sig} \propto \begin{cases} \mathcal{L}^{\rm int} \, \epsilon^2 \, e^{-L_{\rm min}/\bar{d}} & \text{for } \bar{d} \ll L_{\rm min} \\ \mathcal{L}^{\rm int} \, \epsilon^2 \, \frac{L_{\rm max} - L_{\rm min}}{\bar{d}} & \text{for } \bar{d} \gg L_{\rm min} \, . \end{cases}$$

- -- bremsstrahlung
- -- hard-scatterings,...



### NEW PHYSICS FROM PION DECAYS AT THE ATLAS IP

J.L. Feng, I. Galon, F. Kling, ST, 1708.09389

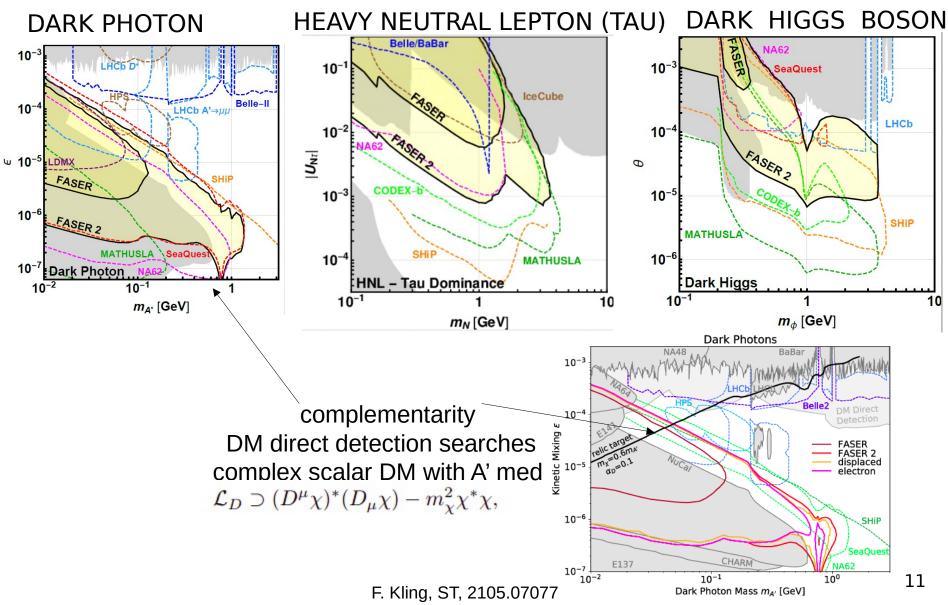


Soft pions going towards high- $p_T$  detectors:

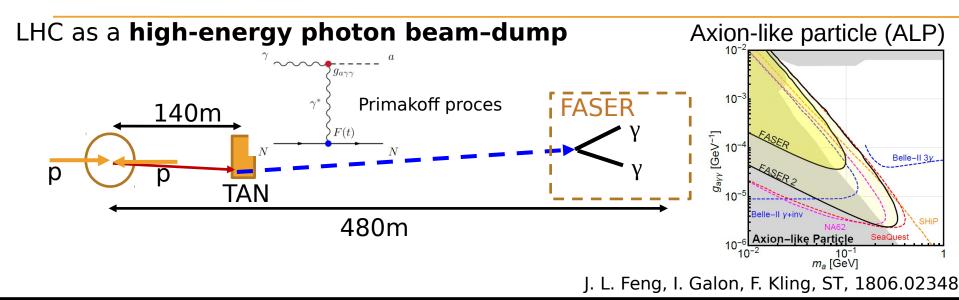
- produced LLPs would be too soft for triggers
- large SM backgrounds

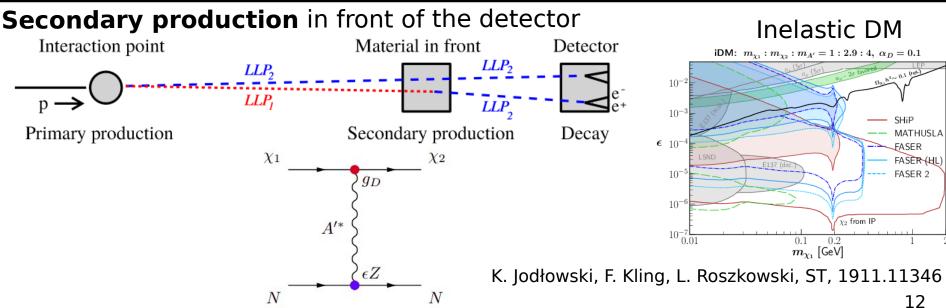
FASER Collaboration, 1811.12522

### SELECTED SENSITIVITY REACH PLOTS



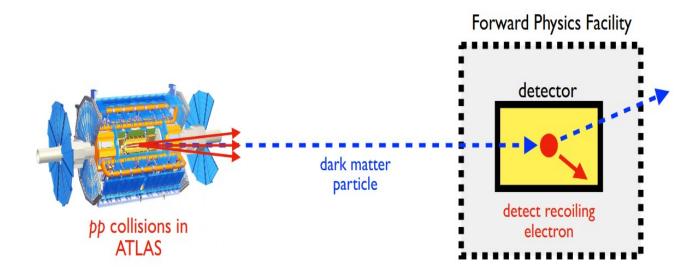
# NOT ONLY ATLAS IP





#### Sebastian Trojanowski (AstroCeNT, CAMK PAN) B. Batell, J. L. Feng, ST, 2101.10338 B. Batell, J.L. Feng, A. Ismail, F. Kling, R.M.Abraham, ST, in preparation DMDIRECT DETECTION AT THE LHC

• Light DM particles can be efficiently produced in the far-forward region of the LHC & scatter in a distance detector



This search is highly complementary to the traditional DM direct detection searches:

 probe of relativistic interaction rates of LDM (DM energy ~ a few hundred GeV) [collider-boosted DM]

– the search is not sensitive to the precise abundance of  $\chi$  DM component (possible variations in cosmological scenario) [collider-produced DM]

## EXAMPLE DM REACH PLOTS

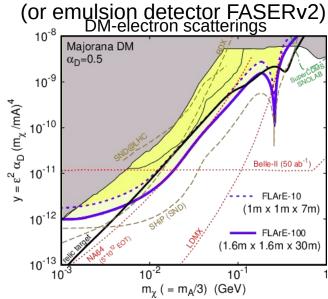
• Useful for probing DM models with suppressed non-relativistic scattering rates

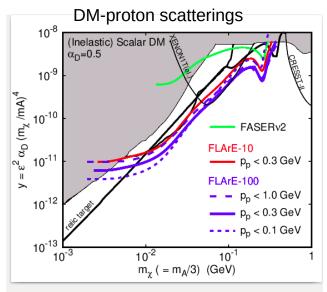
• Sample results for two benchmark models: dark photon mediator & Majorana or (inelastic) complex scalar DM

 $\mathcal{L} \supset A'_{\mu} \left( \epsilon \, e \, J^{\mu}_{EM} + g_D \, J^{\mu}_D \right) \quad \mathcal{L} \supset \begin{cases} |\partial_{\mu}\chi|^2 - m_{\chi}^2 |\chi|^2 & (\text{complex scalar DM}) \\ \frac{1}{2} \overline{\chi} i \gamma^{\mu} \partial_{\mu} \chi - \frac{1}{2} m_{\chi} \overline{\chi} \chi & (\text{Majorana fermion DM}) \end{cases} \qquad J^{\mu}_D = \begin{cases} i \chi^* \overleftrightarrow{\partial_{\mu}} \chi & (\text{complex scalar DM}) \\ \frac{1}{2} \overline{\chi} \gamma^{\mu} \gamma^5 \chi & (\text{Majorana fermion DM}) \end{cases}$ 

They avoid stringent bounds from CMB







### SUMMARY OF FAR-FORWARD BSM PROGRAM

#### (VERY) SCHEMATIC FAR-FORWARD DETECTOR CAPABILITIES Search for LLP decays Scattering detectors: FASERv(2), SND@LHC, FLArE Secondary production FASER(2): 1708.09389, 1811.12522... Current bounds (1911.11346, 2011.04751) 1908.02310. 2001.03073. 2002.08722. 2101.10338 SND@LHC: 2104.09688 DM, v physics, very long-lived new particles Anomalies, DM mediators lifetime (sub-GeV particles) fs μs ps ns

Search for highly-displaced decays of light new particles

(boosted decay lengths d~100-1000 m)

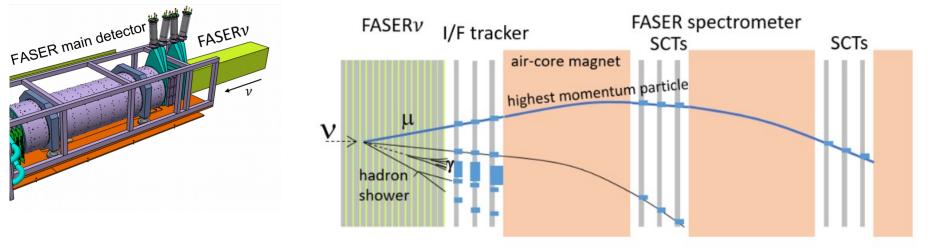
- Secondary production in front of the detector allows for probing even d~meters (or less inside the scattering detectors)
- Scattering detectors:

especially important if decays not possible (stable species like v and DM) can also open new detection channels for very long-lived particles

• Typically best reach for masses<GeV, but even ~10 GeV particles can be probed

# FAR-FORWARD NEUTRINO PROGRAM

### FASERv -- NEUTRINO SUBDETECTOR (RUN 3)



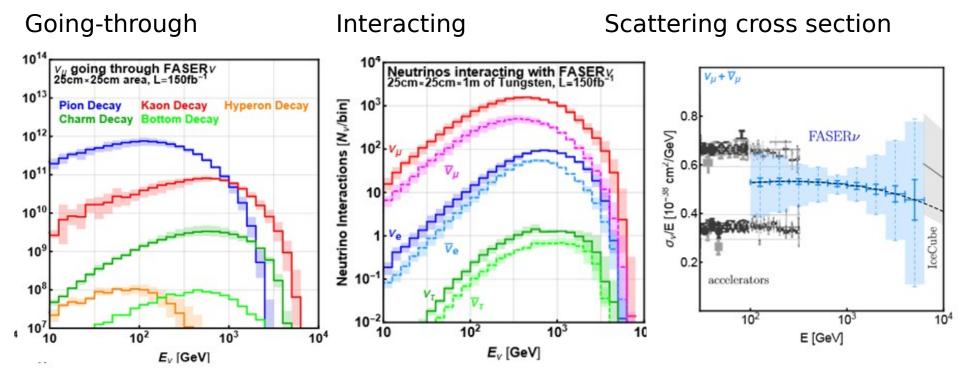
- FASERv (1908.02310, 2001.03073) and <u>SND@LHC</u> (2002.08722) emulsion detectors
- Excellent spatial resolution (even 50nm),
- Can deal with high track density (up to 10<sup>6</sup> tracks/cm<sup>2</sup>),
- Study neutrino interaction vertices at TeV energies in great details
- Interface tracker charge measurement disentangling  $\nu_{\mu}$  and  $\nu_{\mu}$

FASER Collaboration, 1908.02310

# FAR-FORWARD NEUTRINOS

- LHC: lots of forward-going neutrinos from meson decays
- Measurement of the neutrino scattering cross section for  $E_v \sim \text{TeV}$  (currently poorly explored regime)
- Possible detection of 10-20 high-energy tau neutrino events

#### LHC Run 3

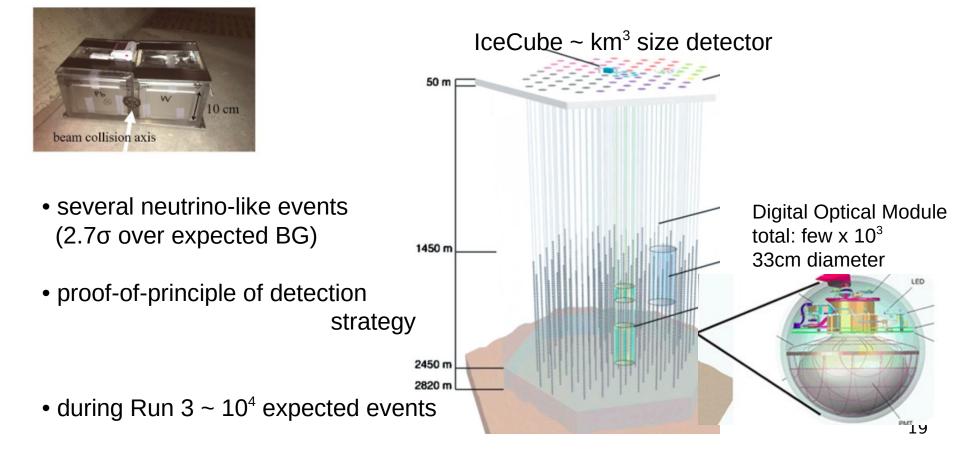


### EXTREMELY POWERFUL DETECTION METHOD

• First neutrino candidate events has been observed already during Run 2...

FASER Collaboration, 2105.06197

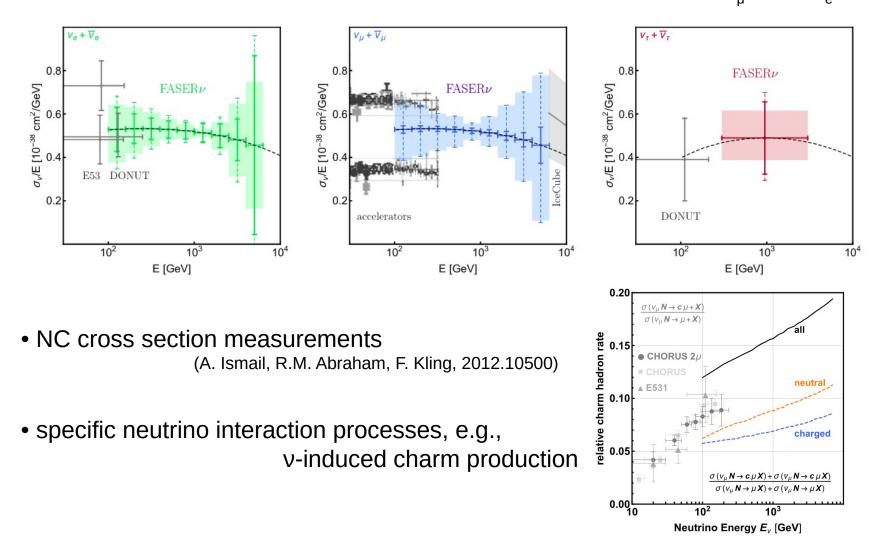
 ...with two handy boxes (10cm x 10cm x 12.5cm) left in the far-forward place (480m) for 4 weeks (12.5 fb<sup>-1</sup>)



FASER Collaboration, 1908.02310

# PROSPECTS FOR RUN 3

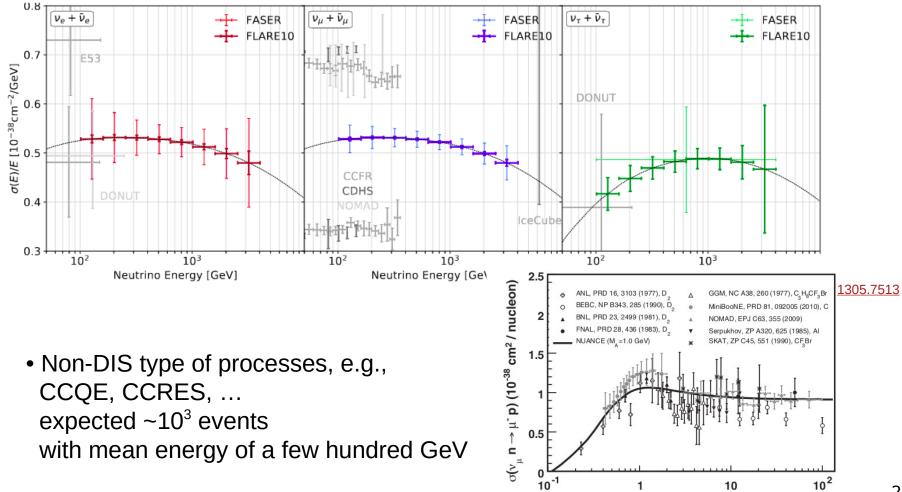
• Inclusive CC cross section measurements at TeV energies (~10<sup>4</sup>  $\nu_{\mu}$ , ~10<sup>3</sup>  $\nu_{e}$ , ~10  $\nu_{r}$ )



E, (GeV)

# PROSPECTS FOR HL-LHC

- 10-tonne detector on beam collision axis
- Even better cross section measurements (few x 10  $^{\scriptscriptstyle 5}\,\nu$  . few x 10  $^{\scriptscriptstyle 4}\,\nu$  . ~10  $^{\scriptscriptstyle 3}\,\nu$  )



### **NEW PHYSICS & NEUTRINO INTERACTIONS**

 Neutrino oscillations into sterile neutrinos direct probes at larger mass differences than typical neutrino experiments

 $\Delta m^2 \sim 1000 \text{ eV}^2$ 

(also e.g. Gallium anomaly)

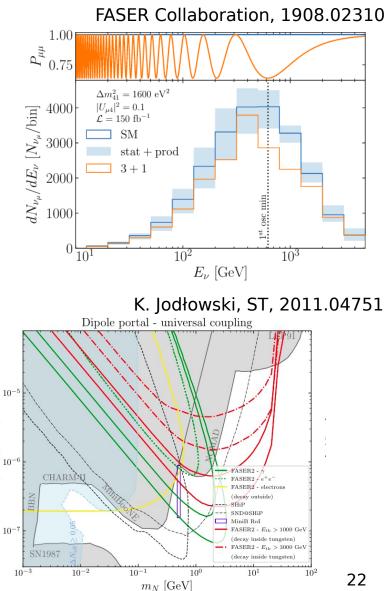
Non-standard neutrino interactions

Example: dipole portal to heavy neutral leptons Magill etal, 1803.03262  $\iota_N \, [1/\text{GeV}]$ 

 $\mathcal{L} \supset \mu_N \, \bar{\nu}_L \sigma_{\mu\nu} N_R F^{\mu\nu} + \text{h.c.},$ 

Transition magnetic moments of neutrinos **Before EWSB** 

$$\mathcal{L} \supset \bar{L} \left( d_{\mathcal{W}} \mathcal{W}^a_{\mu\nu} \tau^a + d_B B_{\mu\nu} \right) \tilde{H} \sigma_{\mu\nu} N_D + h.c.$$

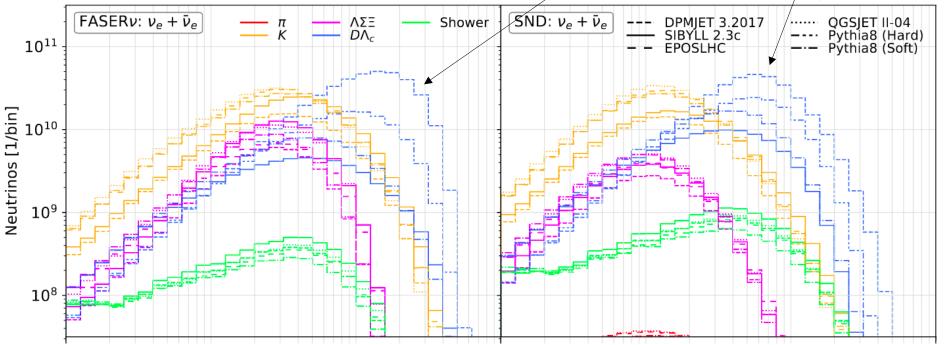


# FURTHER OPPORTUNITIES

## QCD – FORWARD CHARM

- Measuring neutrino flux and spectrum: further tuning of forward MC tools
- Large differences in electron neutrino spectrum at high energies from charm decays
- $v_{a}$  main production at high energies:  $gg \rightarrow cc$ ,  $D \rightarrow K | v$
- probe of gluon PDFs at low x, intrinsic charm,...

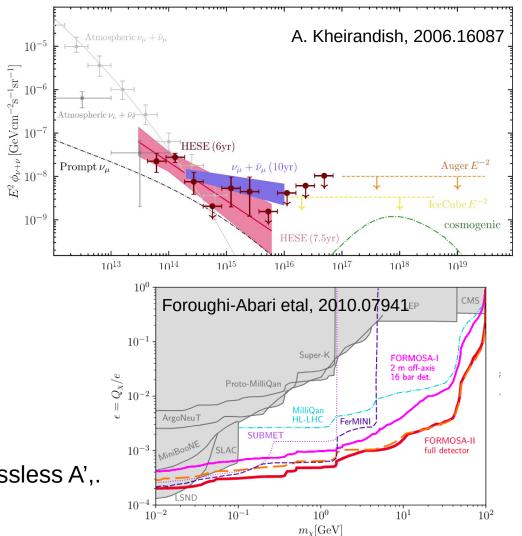
F. Kling, 2105.08270



# COSMIC RAYS AND MORE

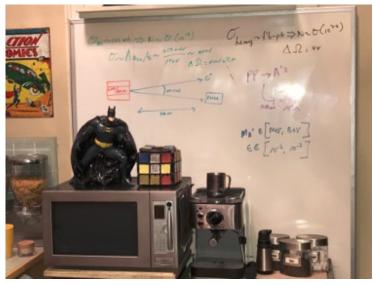
• Forward charm production (relevant for measurements of the astrophysical neutrino flux at IceCube)

- Cosmic-ray muon problem (observed excess of high-energy muons, better high-energy forward kaon production measurement remains essential here)
- Opportunities in muon physics (SM measurements, new physics)
- millicharged particles
- tests of charge quantization
- motivations from GUTs, strings, massless A',.



constrain ``prompt'' atmospheric neutrino flux

# FASER IN POPULAR CULTURE

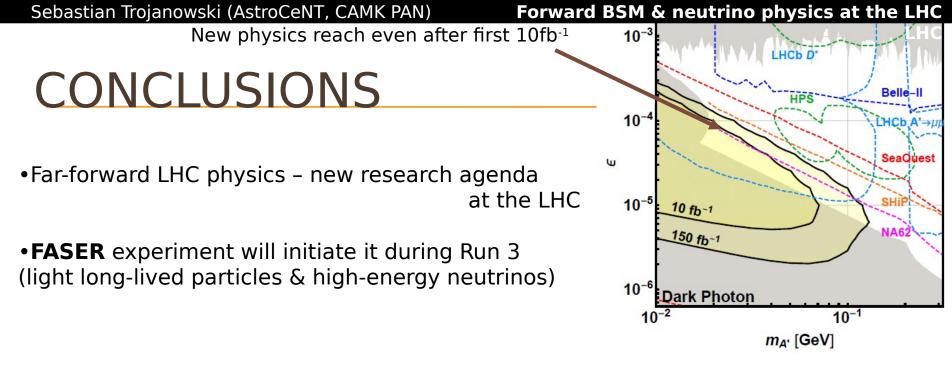






related article





- It could be continued towards HL-LHC: Forward Physics Facility
- further prospects: light DM, QCD and other SM measurements...
- For pheno BSM studies: useful tool **FORESEE** (F. Kling, ST, 2105.07077)
- •2nd Forward Facility Meeting starts just in two days! (>100 participants) https://indico.cern.ch/event/1022352/overview

2nd Forward Physics Facility Meeting

27-28 May 2021